Physics 351 — Friday, March 2, 2018

- Turn in HW6. HW7 will be due the Friday after spring break (3/16), but I haven't written it up yet. I will post it online early next week.
- Read the rest of Ch10 (and answer online questions) during spring break, or as soon as you get back.

Let's try a Lagrangian version of the Foucault pendulum.

 $\mathcal{D} = \hat{z}\cos\theta + \hat{y}\sin\theta$ NL L× x = SLCOSOG- Scina 2 $\int \mathbf{x} \cdot \hat{\mathbf{y}} = - \int \cos \hat{\mathbf{x}}$ X JL x 2 = Jsinox $U = mg L (1 - cos \alpha)$ $\simeq mg L \frac{\alpha^2}{2}$ $\Gamma = X\hat{X} + y\hat{y} + R\hat{z}$ $\Gamma_{0} = \Gamma + \Lambda \times \Gamma = \dot{\chi} \dot{\chi} + \dot{\chi} \dot{\chi} + \Lambda \chi (\dot{\chi} \cos - \hat{z} \sin \theta)$ - RCOSOYX + RSINORX $\vec{r} = (\mathbf{X} - \mathbf{y} \mathcal{R} \cos \theta + \mathbf{R} \mathcal{R} \sin \theta) \hat{\mathbf{X}} + (\hat{\mathbf{y}} + \mathbf{k} \mathcal{R} \cos \theta) \hat{\mathbf{y}} - \mathbf{X} \mathcal{R} \sin \theta \hat{\mathbf{z}}$ $|\Gamma_0|^2 = (x - y \mathcal{R} \cos \theta + R \mathcal{R} \sin \theta)^2 + (y + k \mathcal{R} \sin \theta)^2 + (x \mathcal{R} \sin \theta)^2$ (drop any r2 terms)

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 $\mathbf{F} = (\mathbf{X} - \mathbf{y} \mathcal{R} \cos \theta + \mathbf{R} \mathcal{R} \sin \theta) \hat{\mathbf{X}} + (\mathbf{y} + \mathbf{k} \mathcal{R} \cos \theta) \hat{\mathbf{y}} - \mathbf{X} \mathcal{R} \sin \theta \hat{\mathbf{z}}$ $\left| \int_{0}^{2} \right|^{2} = \left(x - y \mathcal{R} \cos \theta + R \mathcal{R} \sin \theta^{2} + \left(y + k \mathcal{R} \sin \theta^{2} + \left(x \mathcal{R} \sin \theta \right)^{2} + \left(x \mathcal{R} \sin \theta^{2} + \frac{1}{2} \right)^{2} \right)^{2}$ (drop any siz terms) (rol = x-2xy Roso+2x R Rsino +y+2yx Roso $= x^{2} + y^{2} + 2(yk - xy) \operatorname{lcos} + 2xR\operatorname{Rsin} \theta$ $T = \frac{1}{2}m\left[x^{2} + y^{2} + 2(yx - xy) \mathcal{L}\cos\theta + 2xR\mathcal{R}sin\theta\right]$ $U \simeq mq \frac{(\lambda L)^2}{2L} = mg \frac{(\chi^2 + y^2)}{2L} = \frac{mg}{2L} (\chi^2 + y^2)$ $\mathcal{L} = \frac{M}{2} \left[\dot{x}^2 + \dot{y}^2 + 2(\dot{y}x - \dot{x}y) \mathcal{R} \cos + 2\dot{x} \mathcal{R} \mathcal{R} \sin \left[-\frac{M_0^2}{2L} \left(\dot{x}^2 + \dot{y}^2 \right) \right] \right]$

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 $\mathcal{L} = \frac{M}{2} \left[x^{2} + y^{2} + 2(y^{2} - x^{2}y) \mathcal{R}(a \sigma + 2x \mathcal{R} \mathcal{R}(a \sigma)) - \frac{M \sigma}{2L} (x^{2} + y^{2}) \right]$ $\frac{\partial d}{\partial x} = \frac{d}{dt} \left(\frac{\partial d}{\partial x} \right) \Rightarrow my Rox 0 - \frac{mgx}{t} = \frac{d}{dt} \left(\frac{mx}{mx} - my Rox 0 + 2R Rsing \right)$ $\dot{x} - \dot{y} \mathcal{L} \cos \theta = \dot{y} \mathcal{L} \cos \theta - \frac{\partial k}{\mathcal{L}} \implies \dot{x} = -\frac{\partial}{\mathcal{L}} k + 2\mathcal{L} \cos \theta \dot{y}$ $\frac{\partial J}{\partial y} = \frac{1}{dE} \left(\frac{\partial J}{\partial y} \right) \Rightarrow -m_{\chi} \operatorname{Rcoso} - \frac{m_{Q} y}{E} = \frac{1}{dE} \left(m_{\chi} + m_{\chi} \operatorname{Rcoso} \right) = m_{\chi} + m_{\chi} \operatorname{Rcoso}$ $\dot{y} = -\frac{2}{2}y - 2\Lambda \cos x$ $\dot{x} = -\frac{2}{2}x + 2\Lambda \dot{y}$ $bt \exists y = 2\cos y = 2y - 2y \times y$

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$$\begin{aligned} & \text{try } \gamma \equiv x + iy \equiv A e^{iBt} \implies \dot{\eta} \equiv iB\eta, \quad \ddot{\eta} \equiv -B^2\eta \\ & \ddot{\chi} \equiv -\omega_0^2 \chi + 2\lambda_y \dot{y} \qquad - \ddot{y} \equiv -\omega_0^2 \eta - 2\lambda_y \dot{x} \\ & \ddot{\chi} + i\ddot{y} \equiv -\omega_0^2 (\chi + iy) + 2\lambda_y (\dot{y} - i\dot{x}) \equiv -\omega_0^2 (\chi + iy) - 2i\lambda_y (\dot{\chi} + i\dot{y}) \\ & \ddot{\eta} \equiv -\omega_0^2 \eta - 2i\lambda_y \dot{\eta} \implies -B^2 \equiv -\omega_0^2 - (2i\lambda_y)(iB) \\ & -B^2 \equiv -\omega_0^2 + 2\lambda_y B \implies B^2 + 2B\lambda_y - \omega_0^2 \equiv 0 \\ & B \equiv \frac{1}{2} \left(-2\lambda_y \pm \sqrt{4\lambda_y^2 + 4\omega_0^2} \right) \equiv -\lambda_y \pm \sqrt{\lambda_y^2 + \omega_0^2} \implies -\lambda_y \pm \omega_0 \\ & \chi + iy \equiv A_1 e^{-i\lambda_y t} e^{-i\omega_0 t} + A_2 e^{-i\lambda_y t} e^{-i\omega_0 t} \\ & \equiv e^{-i\lambda_y t} \left(C_1 \cos \omega_0 t + C_2 \sin \omega_0 t \right) \\ & \chi(t) \equiv C \cos(\lambda_y t) \cos(\omega_0 t) \\ & \chi(t) \equiv -C \sin(\lambda_y t) \cos(\omega_0 t) \end{aligned}$$

A puck slides with speed v on frictionless ice. The surface is "level" in the sense that it is orthogonal to the effective (gravitational + centrifugal) g at all points. Show that the puck moves in a circle, as seen in Earth's rotating frame. (Assume that v is small enough that the radius of the circle is much smaller than the radius of Earth, so that the colatitude θ is essentially constant throughout the motion.) What is the radius of the circle? What is the frequency of the motion?

Let \hat{x} point east, \hat{y} point north, \hat{z} point "up" so that $\vec{g} = -g\hat{z}$. Earth's rotation vector is \mathcal{R} where $|\mathcal{R}| \simeq \frac{2\pi}{864005} \simeq 7.3 \times 10^{-5} \text{ s}^{-7}$, and $\hat{\gamma}$ eniz + \hat{s} ero = $\hat{\Lambda}$ Forialit = 2m V ×JL = 2m [(Vy lz - Vz ly)x + (Vz lx - Vx lz)y + (Vx ly - Vy lx) 2] Ignore Fz, since we're on frictionless ice, and since we'll assume |Vx Ry | < g, so puck will not go airborne.

 $m_{X} = 2m_{V_{x}}R_{z} = 2m \mathcal{R}\cos V_{y} = m_{V_{x}}$ $m_{y}^{2} = -2m_{V_{x}}R_{z} = -2m \mathcal{R}\cos V_{x} = m_{V_{y}}$ $V_{\chi} = (2R\cos\theta) V_{\chi}$, $V_{\chi} = -(2R)\cos\theta V_{\chi}$ let $\eta = V_x + iV_y = Ae^{i\omega t} \rightarrow \eta = i\omega\eta$ $V_x + iV_y = i\omega V_x - \omega V_x \longrightarrow V_x = -\omega V_y$, $v_y = \omega V_x$ → W=-2 Reaso $X + iy = \frac{A}{i\omega} \left(cos \omega t + is in \omega t \right) = \frac{|A|e^{i\lambda}}{i\omega} \left(cos \omega t + is in \omega t \right)$ If 0=45°, V=15, Han R= 10 km! (Shall effect))

I put this in the Jan 19 notes for reference, but decided it was too tedious to go through in class. In retrospect, I think it's a useful trick to know how to use, so let's do it today.

By the way, there is a fun (and at first glance slightly mysterious) way to prove the dreaded "BAC-CAB rule," using the "Cartesian Einstein notation."

Cartesian Einstein notation
vector
$$\vec{r} = (x_1y_1, z) = x\hat{e}_1 + y\hat{e}_2 + z\hat{e}_3 = \sum r_i \hat{e}_i$$

the it imponent (ie $\{1, 2, 3\}$) of vector \vec{A} is \vec{A}_i
Kronecker delta : $\delta_{ij} = 1$ if $i=j$, else o
dot product: $\vec{A} \cdot \vec{B} = \sum_{ij} A_i B_j \delta_{ij} = \sum_i A_i B_i$
matrix ovector : $\underline{M} \cdot \vec{r} = \sum_{ij} M_{ij} r_j \hat{e}_i$
 $(\underline{M} \cdot \vec{r})_i = \sum_j M_{ij} r_j$
matrix multiply : $(\underline{M} \cdot \underline{M})_{ij} = \sum_k M_{ik} N_{kj}$

http://positron.hep.upenn.edu/p351/files/0119_cartesian_einstein.pdf

Levi-Civita symbol (a.k.c. permutation symbol) antisymmetric symbol) Eijk = S+1 if ijk e (123, 231, 312) -1 if ijk e (213, 321, 132) (D otherwise So $\mathcal{E}_{123} = \mathcal{E}_{231} = \mathcal{E}_{312} = +1$, $\mathcal{E}_{213} = \mathcal{E}_{321} = \mathcal{E}_{132} = -1$ all offens are zero. Cross product: $\vec{A} \times \vec{B} = \sum_{ijk} A_i \vec{B}_j \cdot \vec{e}_k \cdot \vec{E}_{ijk}$ $\vec{A} \times \vec{B} = (A_1 B_2 - A_2 B_1) \hat{e}_3 + (A_2 B_3 - A_3 B_2) \hat{e}_1 + (A_3 B_1 - A_1 B_3) \hat{e}_2$ $(\vec{A} \times \vec{B})_1 = A_2 B_3 - A_3 B_2$ $(\overline{A} \times \overline{B})_2 = (\overline{A}_3 \overline{B}_1 - \overline{A}_1 \overline{B}_3)$ $(\widehat{A} \times \widehat{B})_3 = (A, B_2 - A_2 B_1)$

Incredibly we ful identity: Eijk Elmk = Sie Sim - Sim Sje Now A prove the dreaded "BAC-CAB" rule: $\vec{A} \times (\vec{E} \times \vec{C}) = \sum_{ijk} A_i (\vec{E} \times \vec{C}) \cdot \hat{e}_k \epsilon_{ijk}$ = EA: (Bl Cm en Elmn); ek Eijk = Z A: (Be Cm Elmi) ek Eijk = EA: Be Cm êk (Eemi Eijk)

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 $r = \sum_{ijklim} A_i B_l C_m \hat{e}_k \left(\hat{e}_{lmj} \hat{e}_{ijk} \right)$ $\mathcal{E}_{lmj}\mathcal{E}_{ijk} = \mathcal{E}_{lmj}\mathcal{E}_{kij} = (\mathcal{S}_{lk}\mathcal{S}_{mi} - \mathcal{S}_{li}\mathcal{S}_{mk})$ $\mathcal{V} = \sum_{i \in \mathcal{Q}_{m}} \left(A_{i} B_{\ell} C_{m} \hat{e}_{\kappa} \delta_{\ell \kappa} \delta_{m i} - A_{i} B_{\ell} C_{m} \hat{e}_{\kappa} \delta_{\ell i} \delta_{m \kappa} \right)$ $= \sum \left(A_{i} B_{k} C_{i} \hat{e}_{k} - A_{i} B_{i} C_{k} \hat{e}_{k} \right)$ $= \left(\sum_{i} A_{i}C_{i} \right) \left(\sum_{k} B_{k} \hat{e}_{k} \right) - \left(\sum_{i} A_{i}B_{i} \right) \left(\sum_{k} C_{k} \hat{e}_{k} \right)$ $= (\vec{A} \cdot \vec{c}) \vec{B} - (\vec{A} \cdot \vec{E}) \vec{c}$ $= \vec{B}(\vec{A}\cdot\vec{c}) - \vec{c}(\vec{A}\cdot\vec{B})$

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Let's apply this technique to HW problem 10. To write the kinetic energy w.r.t. the inertial (non-rotating) frame, we use

$$\left(\frac{\mathrm{d} \boldsymbol{r}}{\mathrm{d} t}\right)_{\mathrm{space}} = \left(\frac{\mathrm{d} \boldsymbol{r}}{\mathrm{d} t}\right)_{\mathrm{body}} + \boldsymbol{\Omega} imes \boldsymbol{r}$$

which gives us

$$\boldsymbol{v}_o = \boldsymbol{v} + \Omega imes \boldsymbol{r} = \dot{\boldsymbol{r}} + \Omega imes \boldsymbol{r}$$

where v_o is the velocity in the inertial ("space") frame, while v is the velocity in the rotating ("body") frame. Now we can use v_o to write the KE w.r.t. the inertial frame:

$$\mathcal{L} = \frac{1}{2}m|\boldsymbol{v}_o|^2 - U(\boldsymbol{r}) = \frac{1}{2}m|\dot{\boldsymbol{r}} + \boldsymbol{\Omega} \times \boldsymbol{r}|^2 - U(\boldsymbol{r})$$

Writing out the KE component by component (i=x,y,z):

$$\mathcal{L} = \left[\sum_{i} \frac{m}{2} (\dot{\boldsymbol{r}} + \boldsymbol{\Omega} \times \boldsymbol{r})_{i}^{2}\right] - U(\boldsymbol{r})$$

Now pick coordinate n and differentiate \mathcal{L} . As usual, $\partial A^2/\partial r_n = (2A)(\partial A/\partial r_n)$.

$$\frac{\partial \mathcal{L}}{\partial r_n} = \left[\sum_i m(\dot{\boldsymbol{r}} + \boldsymbol{\Omega} \times \boldsymbol{r})_i \frac{\partial}{\partial r_n} (\dot{r}_i + (\boldsymbol{\Omega} \times \boldsymbol{r})_i)\right] - \frac{\partial U(\boldsymbol{r})}{\partial r_n}$$

where I rewrote $(\dot{\mathbf{r}} + \mathbf{\Omega} \times \mathbf{r})_i$ as $\dot{r}_i + (\mathbf{\Omega} \times \mathbf{r})_i$. The derivative of the first term is zero: $\partial \dot{r}_i / \partial r_n = 0$. We can write out the second term using the Cartesian Einstein notation as

$$(\mathbf{\Omega} \times \mathbf{r})_i = \sum_{jk} \Omega_j r_k \epsilon_{ijk}$$

whose derivative is

$$\frac{\partial}{\partial r_n} (\mathbf{\Omega} \times \mathbf{r})_i = \sum_{jk} \Omega_j \left(\frac{\partial r_k}{\partial r_n} \right) \epsilon_{ijk} = \sum_{jk} \Omega_j \delta_{kn} \epsilon_{ijk} = \sum_j \Omega_j \epsilon_{ijn}$$

Now we can plug this in to $\partial \mathcal{L} / \partial r_n$

$$\frac{\partial \mathcal{L}}{\partial r_n} = \left[\sum_i m(\dot{\boldsymbol{r}} + \boldsymbol{\Omega} \times \boldsymbol{r})_i \left(\sum_j \Omega_j \epsilon_{ijn}\right)\right] - (\boldsymbol{\nabla} U)_n$$
$$\frac{\partial \mathcal{L}}{\partial r_n} = \left[\sum_{ij} m(\dot{\boldsymbol{r}} + \boldsymbol{\Omega} \times \boldsymbol{r})_i \Omega_j \epsilon_{ijn}\right] - (\boldsymbol{\nabla} U)_n$$

Then using $\sum_{ij} A_i B_j \epsilon_{ijn} = (\mathbf{A} \times \mathbf{B})_n$ we rewrite this as a cross-product:

$$\frac{\partial \mathcal{L}}{\partial r_n} = m \left[(\dot{\boldsymbol{r}} + \boldsymbol{\Omega} \times \boldsymbol{r}) \times \boldsymbol{\Omega} \right]_n - (\boldsymbol{\nabla} U)_n$$

and then distribute

$$\frac{\partial \mathcal{L}}{\partial r_n} = m(\boldsymbol{v} \times \Omega)_n + m((\boldsymbol{\Omega} \times \boldsymbol{r}) \times \Omega)_n - (\boldsymbol{\nabla} U)_n$$

Now let's go back and differentiate \mathcal{L} w.r.t. \dot{r}_n (dropping the potential term since $\partial U/\partial \dot{r}_n = 0$)

$$\frac{\partial \mathcal{L}}{\partial \dot{r}_n} = \sum_i m(\dot{\boldsymbol{r}} + \boldsymbol{\Omega} \times \boldsymbol{r})_i \frac{\partial}{\partial \dot{r}_n} (\dot{r}_i + (\boldsymbol{\Omega} \times \boldsymbol{r})_i)$$

Then use $\partial \dot{r}_i / \partial \dot{r}_n = \delta_{in}$ and $\partial r_i / \partial \dot{r}_n = 0$

$$\frac{\partial \mathcal{L}}{\partial \dot{r}_n} = \sum_i m(\dot{\boldsymbol{r}} + \boldsymbol{\Omega} \times \boldsymbol{r})_i \, \delta_{in} = m(\dot{\boldsymbol{r}} + \boldsymbol{\Omega} \times \boldsymbol{r})_n = mv_n + m(\boldsymbol{\Omega} \times \boldsymbol{r})_n$$

Now take the time derivative:

$$\frac{\mathrm{d}}{\mathrm{d}t}\frac{\partial \mathcal{L}}{\partial \dot{r}_n} = ma_n + m(\dot{\boldsymbol{\Omega}} \times \boldsymbol{r})_n + m(\boldsymbol{\Omega} \times \boldsymbol{v})_n$$

So the Lagrange equation of motion for component r_n reads

$$ma_n + m(\dot{\boldsymbol{\Omega}} \times \boldsymbol{r})_n + m(\boldsymbol{\Omega} \times \boldsymbol{v})_n = m(\boldsymbol{v} \times \Omega)_n + m((\boldsymbol{\Omega} \times \boldsymbol{r}) \times \boldsymbol{\Omega})_n - (\boldsymbol{\nabla} U)_n$$

Combining the components into vectors,

$$m\boldsymbol{a} + m\dot{\boldsymbol{\Omega}} \times \boldsymbol{r} + m\boldsymbol{\Omega} \times \boldsymbol{v} = m\boldsymbol{v} \times \boldsymbol{\Omega} + m(\boldsymbol{\Omega} \times \boldsymbol{r}) \times \boldsymbol{\Omega} - \boldsymbol{\nabla} U$$

Then permute to flip signs, and use $F = -\nabla U$, we have $U = -\nabla Q$.

$$m\boldsymbol{a} - m\boldsymbol{r} \times \dot{\boldsymbol{\Omega}} - m\boldsymbol{v} \times \boldsymbol{\Omega} = m\boldsymbol{v} \times \boldsymbol{\Omega} + m(\boldsymbol{\Omega} \times \boldsymbol{r}) \times \boldsymbol{\Omega} + \boldsymbol{F}$$

and rearrange to get vector sum of real force and the three pseudo-forces: "azimuthal" (a.k.a. Euler) force, Coriolis force, and centrifugal force.

$$m\boldsymbol{a} = \boldsymbol{F} + m\,\boldsymbol{r} \times \dot{\boldsymbol{\Omega}} + 2m\,\boldsymbol{v} \times \boldsymbol{\Omega} + m\,(\boldsymbol{\Omega} \times \boldsymbol{r}) \times \boldsymbol{\Omega}$$

http://positron.hep.upenn.edu/p351/files/0302_pseudoforce.pdf

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Enjoy your week off. Safe travels!