# Physics 351 — Wednesday, March 28, 2018

- I have graded 40/60 of the midterm exams. I'll finish tomorrow and will hand them back Friday.
- HW9 due Friday. For HW help, Bill is in DRL 3N6 Wed 4–7pm. Grace is in DRL 2C2 Thu 5:30–8:30pm. To get the most benefit from the homework, first work through every problem on your own to the best of your ability. Then check in with me, Grace, or a friend to compare final results and to trade suggestions on problems that stumped you.
- For the next 3–4 days, we will finally make our way through the rest of Chapter 10, which is (I think) the most difficult chapter in the book. The two significant topics after that are coupled oscillators and the Hamiltonian formalism, plus a bit of "enrichment" material.
- ► FYI intuitive description of precession:

http://positron.hep.upenn.edu/p351/files/0331\_george\_abell\_precession.pdf

## Morin Exerise 9.38.

### 9.38. Striking a triangle \*\*

Consider the rigid object in Fig. 9.57. Four masses lie at the points shown on a rigid isosceles right triangle with hypotenuse length 4a. The mass at the right angle is 3m, and the other three masses are m. Label them A, B, C, D, as shown. Assume that the object is floating freely in outer space. Mass C is struck with a quick blow, directed into the page. Let the impulse have magnitude  $\int F dt = P$ . What are the velocities of all the masses immediately after the blow? 2 sut of pure



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3m 2a20 m m Fig. 9.57

Where is the CM? Let's call the CM (initially) (0, 0, 0). What is the post-impact motion of the CM? What are the principal axes/moments? Write two different expressions for L, to find  $\omega$ . Use  $\omega$  to find velocities w.r.t. CM, then combine with  $v_{
m CM}$ .

## Morin Exerise 9.38.

### 9.38. Striking a triangle \*\*

Consider the rigid object in Fig. 9.57. Four masses lie at the points shown on a rigid isosceles right triangle with hypotenuse length 4a. The mass at the right angle is 3m, and the other three masses are m. Label them A, B, C, D, as shown. Assume that the object is floating freely in outer space. Mass C is struck with a quick blow, directed into the page. Let the impulse have magnitude  $\int F dt = P$ . What are the velocities of all the masses immediately after the blow?  $\Im$ 



FIG. 9.50

Where is the CM? Let's call the CM (initially) (0,0,0). Halfway between B and D.

What is the post-impact motion of the CM?

 $V_{\rm cm} = -P/(6m)\hat{z}$  (and stays that way) What are the principal axes/moments (w.r.t. CM)?

 $\lambda_1 = 6ma^2 \ (\hat{\boldsymbol{e}}_1 = \hat{\boldsymbol{x}}), \ \lambda_2 = 8ma^2 \ (\hat{\boldsymbol{e}}_2 = \hat{\boldsymbol{y}}), \ \lambda_3 = 14ma^2 \ (\hat{\boldsymbol{e}}_3 = \hat{\boldsymbol{z}})$ Write two different expressions for  $\boldsymbol{L}$ , to find  $\boldsymbol{\omega}$ . Use  $\boldsymbol{\omega}$  to find velocities w.r.t. CM, then combine with  $\boldsymbol{v}_{\rm CM}$ .

$$\begin{array}{c} D=3m_{\odot}(0,fa,0)\\ put arigin at CMA \\ f=m_{\odot}(2a,-a,0)\\ f=m_{\odot}(2a,-a,-b,-a,-b,-a,-b)\\ f=m_{\odot}(2a,-a,-b)\\ f=m_{\odot}(2a,-a,-b)\\ f=m_{\odot}(2a,-a,-b)\\ f=m_{\odot}(2a,-a,-b)\\ f=m_{\odot}(2a,-a,-b)\\ f=m_{\odot}(2a,-a,-b)\\ f=m_{\odot}(2a,-b)\\ f=m_{\odot}(2a,-b)$$

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A B C  $\vec{V}_{introdictuly} = \vec{V}_{cm} + \vec{\omega} \times \vec{r}$ after import  $\vec{V}_{cm} = \frac{P}{m}(0, 0, -\frac{1}{6})$  $\vec{W} \times \vec{F}_{A} = \frac{P}{ma} (\frac{1}{6}, \frac{1}{7}, 0) \times (-2a, -a, 0) = \frac{P}{m} (0, 0, \frac{1}{3})$  $\vec{\omega} \times \vec{r}_{B} = \frac{P}{ma} \left( \frac{1}{6}, \frac{1}{7}, 0 \right) \times (0, -\alpha, 0) = \frac{P}{m} \left( 0, 0, -\frac{1}{6} \right)$  $\vec{\omega} \times \vec{r}_{c} = \frac{P}{ma} \left( \frac{1}{6}, \frac{1}{7}, 0 \right) \times (+2a, -a, 0) = \frac{P}{m} \left( 0, 0, -\frac{2}{3} \right)$  $\vec{u} \times \vec{r}_{D} = \frac{P}{ma} \left( \frac{1}{6}, \frac{1}{7}, 0 \right) \times \left( 0, +a, 0 \right) = \frac{P}{m} \left( 0, 0, \frac{1}{6} \right)$  $\implies \vec{V}_{A} = \frac{P}{m}(0, 0, \frac{1}{6}) / \vec{V}_{c} = \frac{P}{m}(0, 0, -\frac{5}{6})$  $\vec{V}_{B} = \frac{P}{M} \left( 0, 0, -\frac{1}{3} \right) \quad \vec{V}_{D} = \frac{P}{M} \left( 0, 0, 0 \right)$ 

In[9]:=

$$\{1/6, 1/4, 0\} \times \{-2, -1, 0\}$$
  
Out[9]=  $\{0, 0, \frac{1}{3}\}$ 

 $ln[10]:= \{1/6, 1/4, 0\} \times \{0, -1, 0\}$ Out[10]=

$$\left[0, 0, -\frac{1}{6}\right]$$

 $\ln[11]:= \{1/6, 1/4, 0\} \times \{+2, -1, 0\}$ 

Out[11]=

$$\left\{0, 0, -\frac{2}{3}\right\}$$

 $ln[12]:= \{1/6, 1/4, 0\} \times \{0, +1, 0\}$ 

Out[12]=

$$\left\{0, 0, \frac{1}{6}\right\}$$

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What will the subsequently happen to  $V_{\rm cm}$ ? To L? To  $\omega$ ? To the orientations of the principal axes? With no applied torque, how does  $\omega$  evolve in time?





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> https://www.youtube.com/watch?v=dVhGyxkBKzI https://www.youtube.com/watch?v=4Ntgvun8GuY https://www.youtube.com/watch?v=YKSEu\_c3YdY

ero torque)  $=\left(\frac{dL}{dt}\right)_{Space}$ = (de) + wxL  $(\omega, \lambda_1, \omega, \lambda_2, \omega, \lambda_2) = -\omega \times (\omega, \lambda_1, \omega, \lambda_2, \omega, \lambda_2)$  $(\omega_2\omega_3\lambda_3-\omega_3\omega_2\lambda_2,\omega_3\omega_1\lambda_1-\omega_1\omega_3\lambda_3,\omega_1\omega_2\lambda_2-\omega_1\omega_3\lambda_3)$ W2W, L  $(\omega_2\omega_3(\lambda_2-\lambda_3),\omega_1\omega_3(\lambda_3-\lambda_1),\omega_1\omega_2(\lambda_1-\lambda_2))$  $\omega_{1} = \omega_{2}\omega_{3} \frac{(\lambda_{2} - \lambda_{3})}{\lambda_{1}}$  $\omega_1 = \omega_1 \omega_2 \frac{\lambda_1 - \lambda_2}{\omega_1 - \omega_2}$ w.  $= \omega_1 \omega_2 \lambda_3 - \lambda_1$ 

It's fun to consider e.g.  $\lambda_3 > \lambda_2 > \lambda_1$  for tossed book.

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Start out e.g. about êz, Wza wi  $W = W_2 W_3 \frac{\lambda_2 - \lambda_3}{\lambda}$  $\omega_z = \omega_1 \omega_2$  $W_1 \simeq W_2 W_3 \frac{\lambda_2 - \lambda_3}{\lambda_1}$ 'Small x Small => W3 ~ constant  $\omega_2 = \omega_1 \omega_3 \frac{\gamma_3 - \gamma_1}{\gamma_1}$  $\mathcal{Y} \stackrel{\sim}{\omega} = \omega, \left(\omega_3 \frac{\lambda_3 - \lambda_1}{\lambda_2}\right) \left(\omega_3 \frac{\lambda_2 - \lambda_3}{\lambda_2}\right) = -\omega,$ W3  $\omega_2 \simeq \omega, \ \omega_3 \frac{\lambda_3 - \lambda_1}{\lambda_2} = \omega_2 (\omega_3 \frac{\lambda_2 - \lambda_3}{\lambda_1}) (\omega_3 \frac{\lambda_3 - \lambda_1}{\lambda_2})$  $\omega_2 \simeq -\omega_2 \left( \omega_3^2 \left( \frac{\lambda_3 - \lambda_2}{\lambda_3 - \lambda_2} \right) \right) =$ -REW2

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If you look down the  $\hat{e}_3$  axis, you'll see the tip of  $\omega$  tracing out an ellipse whose ratio of axis lengths is  $\sqrt{\frac{(\lambda_3 - \lambda_2)\lambda_2}{(\lambda_3 - \lambda_1)\lambda_1}}$ .

$$\begin{split} \lambda_{i}\dot{\omega_{1}} + \left[ \left( \omega_{i}, \omega_{2}, \omega_{3} \right) \times \left( \lambda_{i}\omega_{i}, \lambda_{2}\omega_{2}, \lambda_{3}\omega_{3} \right) \right]_{i} = 0 \\ \lambda_{i}\dot{\omega_{1}} + \left( \omega_{2}\lambda_{3}\omega_{3} - \omega_{3}\lambda_{2}\omega_{2} \right) = 0 \\ \lambda_{i}\dot{\omega_{1}} + \omega_{2}\omega_{3}\left( \lambda_{3} - \lambda_{2} \right) = 0 \\ \dot{\omega_{1}} = \omega_{2}\omega_{3}\frac{\lambda_{2} - \lambda_{3}}{\lambda_{1}} \rightarrow \ddot{\omega_{1}} = \dot{\omega}_{2}\left( \omega_{3}\frac{\lambda_{2} - \lambda_{3}}{\lambda_{1}} \right) = \omega_{i}\left( \omega_{3}\frac{\lambda_{2} - \lambda_{3}}{\lambda_{1}} \right) \\ \dot{\omega_{2}} = \omega_{3}\omega_{i}\frac{\lambda_{3} - \lambda_{1}}{\lambda_{2}} \rightarrow \ddot{\omega_{1}} = \dot{\omega}_{1}\left( \omega_{3}\frac{\lambda_{3} - \lambda_{1}}{\lambda_{2}} \right) = \omega_{2}\left( \omega_{3}^{2}\frac{\left( \lambda_{2} - \lambda_{3} \right)\left( \lambda_{3} - \lambda_{1} \right)}{\lambda_{2}} \right) \\ \omega_{1} = A\cos\Omega t \rightarrow \dot{\omega_{1}} = -\Omega A\sin\Omega t = -\frac{\Omega A}{B}\omega_{2} \rightarrow \omega_{3}\frac{\lambda_{3} - \lambda_{2}}{\lambda_{1}} = \frac{\Omega B}{B} \\ \omega_{2} = B\sin\Omega t \rightarrow \dot{\omega_{2}} = \Omega B\cos\Omega t = \frac{\Omega B}{A}\omega_{1} \rightarrow \omega_{3}\frac{\lambda_{3} - \lambda_{2}}{\lambda_{1}} = \frac{\Omega B}{A} \\ \frac{A^{2}}{B^{2}} = \frac{\left( \lambda_{3} - \lambda_{2} \right)\lambda_{2}}{\left( \lambda_{3} - \lambda_{1} \right)\lambda_{1}} \end{split}$$

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 $\lambda_2 < \lambda_3), s_0$ Start out about e W2 and W2 both W2 and < W, initially.  $\omega = \omega_2 \omega_3 \frac{\lambda_2 - \lambda_3}{1}$ N.  $= W_1 W_3 \frac{\lambda_3}{\sqrt{2}}$  $\sim |Small)^2$ w2 13-~ W. Wz 2-11) K3  $= \omega_1 \omega_2 \frac{\lambda_1 - \lambda_2}{\lambda_2}$ R<sup>2</sup>W3 W1 13 11/212 WZ ~ W, WZ -W2

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Start out about ez, so w, and wz « wz initially.  $\omega_2 = \omega_1 \omega_3 \frac{\lambda_3 - \lambda_1}{\lambda} \sim (Small)^2$  $\omega_1 = \omega_2 \omega_3 \frac{\lambda_2 - \lambda_3}{\lambda_2}$  $\overset{\circ\circ}{W}_{1} \simeq W_{2} \left( \underbrace{W_{1}}_{\lambda_{2}} \underbrace{\lambda_{1} - \lambda_{2}}_{\lambda_{2}} \right) \underbrace{\lambda_{2} - \lambda_{3}}_{\lambda_{1}} = + \underbrace{W_{1}}_{\lambda_{1}} \left( \underbrace{W_{2}}_{2} \underbrace{\lambda_{2} - \lambda_{1}}_{\lambda_{1}} \right) \underbrace{\lambda_{2} - \lambda_{2}}_{\lambda_{1}} \right)$  $\dot{\omega}_{3} \simeq \left(\omega_{2}\omega_{3} \quad \frac{\lambda_{2}-\lambda_{3}}{\lambda_{1}}\right)\omega_{2} \quad \frac{\lambda_{1}-\lambda_{2}}{\lambda_{3}} = +\omega_{3}\left(\omega_{2} \frac{(\lambda_{2}-\lambda_{1})(\lambda_{3}-\lambda_{2})}{\lambda_{1}\lambda_{3}}\right)$ => exponential growth of winwz -> initial motion about ez won't stay about

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e and body axes coincide at 
$$t = 0$$
  
 $= \frac{P}{ma}(\frac{1}{6}, \frac{1}{4}, 0)$ .  $L \equiv aP(1, 2, 0)$ .  
 $\dot{\omega}_1 = \omega_2 \omega_3 \frac{\lambda_2 - \lambda_3}{\lambda_1}$   
 $\dot{\omega}_2 = \omega_3 \omega_1 \frac{\lambda_3 - \lambda_1}{\lambda_2}$   
 $\dot{\omega}_3 = \omega_1 \omega_2 \frac{\lambda_1 - \lambda_2}{\lambda_3}$ 

 $\label{eq:http://positron.hep.upenn.edu/p351/files/0327_strucktriangle.nb http://positron.hep.upenn.edu/p351/files/0327_strucktriangle.pdf http://positron.hep.upenn.edu/p351/files/0327_strucktriangle_230.avi https://www.youtube.com/watch?v=IMBRIyxDLss Consider how you would go about calculating the <math>(x, y, z)$  (space) positions of vertices A, C, D vs. time. I did it by keeping track of the (x, y, z) coordinates of the unit vectors  $\hat{e}_1, \hat{e}_2, \hat{e}_3$  as a function of time.

(Neglecting Ven, as I did in animation) +292, -982 aez ·Zaê 5 2 9 8 Ven Dod in 0 1954 Space Space UDdat 10 me updat body. Ima

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 $\Gamma = \left(\frac{dL}{dt}\right)_{\text{space}} = \left(\frac{dS}{dt}\right)_{\text{body}} + \omega \times L$  $(\Gamma, \Gamma_2, \Gamma_3) = (\lambda, \omega_1, \lambda_2 \omega_2, \lambda_3 \omega_3) + \omega \times \zeta$  $\Gamma_3 = \lambda_3 \omega_3 + (\omega_1 L_2 - \omega_2 L_1)$  $\Gamma_3 = \lambda_2 \omega_2 + \omega_1 \omega_2 \lambda_2 - \omega_2 \omega_1 \lambda_1$  $\Gamma_3 = \lambda_2 \omega_2 + (\lambda_2 - \lambda_1) \omega_1 \omega_2$  $\lambda_3 \upsilon_2 = \Gamma_3 + (\lambda_1 - \lambda_2) \omega_1 \omega_2$  $\lambda, \omega = \Gamma, + (\lambda_2 - \lambda_3) \omega_2 \omega_3$  $\lambda_2 \omega_2 = \Gamma_2 + (\lambda_3 - \lambda_1) \omega_2 \omega_1$ 

Torque-free precession of symmetric top (more on this later):

Hen X, w, = 13 (Ix=2)  $I \in \lambda, = \lambda = \lambda$ If  $\Gamma_2 = 0$  then  $\omega_3 = const.$  Suppose  $\Gamma = 0$ . -13) W2 W2 let . ) wzwy - W,  $w = - \mathcal{L} w,$ = w, cosset 12. No =- No shilt

As seen from body frame,  $\omega$  precesses about  $\hat{e}_3$  with frequency  $\Omega$ . As seen from the body frame, what does L do?

What does the situation look like from the space frame?

If  $\lambda_1 = \lambda_2 = \lambda$  then  $\lambda_2 \omega_2 = 1_3$  $I_{X,=Z_{r}}$ If I3=0 then w3 = const. Suppose I = 0.  $\lambda \omega_1 = P_1 + (\lambda - \lambda_2) \omega_2 \omega_3$ let . 2/  $\dot{w}_2 = P_2 + (\lambda_3 - \lambda) w_2 w_1$  $W_1 = \mathcal{L} W_2$ ,  $W_2 = -\mathcal{L} W_2$ -> W = W cosset 12=-No shilt

As seen from body frame, L and  $\omega$  precess about (fixed)  $\hat{e}_3$  with frequency  $\Omega_b \equiv \Omega = \omega_3(\lambda - \lambda_3)/\lambda$ , where  $\lambda = \lambda_1 = \lambda_2$ .

As seen from the space frame,  $\hat{e}_3$  and  $\omega$  precess about (fixed) L, at a frequency that takes some effort to calculate. (You'll calculate the space-frame precession frequency,  $\Omega_s$ , on a HW problem next week. It is much more involved than you might expect.)

Video from two 2015 students traveling back from spring break: https://www.youtube.com/watch?v=bVpPp1e\_1Z4

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Astronaut version:
https://youtu.be/fPI-rSwAQNg
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Cosmonaut version (!): Dzhanibekov effect 
https://youtu.be/dL6Pt10_gSE
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https://www.youtube.com/watch?v=BGRWg4aV2mw

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Someone's quasi-intuitive explanation:
http://mathoverflow.net/questions/81960/
the-dzhanibekov-effect-an-exercise-in-mechanics-or-fiction-explain-mathemat
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(Taylor 10.35) A rigid body consists of: m at (a, 0, 0) = a(1, 0, 0) 2m at (0, a, a) = a(0, 1, 1) 3m at (0, a, -a) = a(0, 1, -1)Find inertia tensor  $\underline{I}$ , its principal moments, and the principal axes.

 $I_{kk} = Z_{M}(y^{2}+z^{2}) = Ma^{2}(Z^{2}+2^{3}) = 10ma^{2}$  $I_{yy} = Z_m(x^2+z^2) = Ma^2(1+z+3) = 6ma^2$  $I_{22} = Zm(x^2+y^2) = mq^2(1+2+3) = (emq^2)$  $I_{XY} = -\Sigma m_{XY} = -ma^2(0) = 0$  $I_{XZ} = - Z_{MXZ} = -MG^{2}(0) = 0$  $I_{yz} = -\Sigma my_{z} = -ma^{2}(2-3) = ma^{2}$  $\begin{aligned}
\Xi &= \begin{pmatrix} 10 & 0 & 0 \\ 0 & 6 & 1 \end{pmatrix} Ma^2 \\
& & & & & \\ 0 & 1 & 6 \end{pmatrix}
\end{aligned}$ ▲ロト ▲圖 ト ▲ 国 ト ▲ 国 ト ● ④ ● ●

 $\omega = \lambda \omega \implies (\underline{I} \rightarrow \lambda)$  $\det (\underline{T} - \lambda \underline{1}) = 0$  $0 = (10 - \lambda)(6 - \lambda)^{2} - (10 - \lambda) \Rightarrow \lambda = 10 \text{ or } (6 - \lambda)^{2} = 1$   $b - \lambda = 1 \Rightarrow \lambda = 5, \quad 6 - \lambda = -1 \Rightarrow \lambda = 7 \quad \lambda \in \{10, 7, 5\}$ 

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57 maz 10, 10-10 0 0 0 3 D = 6-10 С = ۵ -4 6-10 0 0 00 4 0 シュ ΰ 6-y2 0 -2 0 0 0 2 0 10 33 -0 6-5 0 5



### eigenvectors {{10,0,0},{0,6,1},{0,1,6}}

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 $In[1]:= m = \{\{10, 0, 0\}, \{0, 6, 1\}, \{0, 1, 6\}\}$  $Dut[1]= \{\{10, 0, 0\}, \{0, 6, 1\}, \{0, 1, 6\}\}$ 

In[2]:= MatrixForm[m]

2]//MatrixForm=

 $\left(\begin{array}{rrrrr}
10 & 0 & 0\\
0 & 6 & 1\\
0 & 1 & 6
\end{array}\right)$ 

In[3]:= Eigenvalues[m]
Dut[3]= {10, 7, 5}

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One useful tool for relating the fixed  $\hat{x}$ ,  $\hat{y}$ ,  $\hat{z}$  axes to the rigid body's  $\hat{e}_1$ ,  $\hat{e}_2$ ,  $\hat{e}_3$ , axes is the "Euler angles,"  $\phi$ ,  $\theta$ ,  $\psi$ .

(Another way, which I used in the simulation program for the struck triangle, is simply to keep track instant-by-instant of the x, y, z components of  $\hat{e}_1(t)$ ,  $\hat{e}_2(t)$ ,  $\hat{e}_3(t)$ . But if you're given the three Euler angles, you can compute the x, y, z components of the body axes  $\hat{e}_1$ ,  $\hat{e}_2$ ,  $\hat{e}_3$ .)

Question: Suppose I rotate the vector  $(x, y) = R(\cos \alpha, \sin \alpha)$  by an angle  $\phi$  (about the origin). How would you write x' as a linear combination of x and y? How about y' as a linear combination of x and y?



Rotate by angle & about 2  $\begin{pmatrix} x' \\ y' \end{pmatrix} = \begin{pmatrix} cos\phi & -sin\phi \\ sin\phi & cos\phi \end{pmatrix}$  $\chi' = \chi \cos \phi - \gamma \sin \chi$  $y' = \chi sin\phi + \gamma cos\phi$ R (Ky) (X, y)=R(COSK, Sind)  $(X, y') = R(\cos(\alpha + \phi), \sin(\alpha + \phi))$ = R (cosk cosk - sinksing, sind cosk + cosksing) = (xcorp-ysing, ycorp + ksing)

Rotate angle \$ about Iniz-(ask Sind 0 l'as above ! 2'=2 angle & about 0200 Sino Case

Mnemonic: for infinitessimal rotation angle  $\epsilon \ll 1$ ,  $r \to r + \epsilon \hat{\omega} \times r$ . So for rotation about  $\hat{y}$ ,  $(1,0,0) \to (1,0,-\epsilon)$ , since  $\epsilon \hat{y} \times \hat{x} = -\epsilon \hat{z}$ .

The hardest part of writing down  $3 \times 3$  rotation matrices is remembering where to put the minus sign.



Once you've worked out one case correctly (e.g. from a diagram), here's a trick (thanks to 2015 student Adam Zachar) for working out the other two ...

Just add two more columns and two more rows, following the cycles: xyz, yzx, zxy. Then draw boxes of size  $3 \times 3$ .



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(Check previous result using Mathematica.)

In[1]:= RotationMatrix[\$\phi\$, {0, 0, 1}] // MatrixForm

Out[1]//MatrixForm=

$$egin{array}{c} \cos \left[ \phi 
ight] & -\sin \left[ \phi 
ight] & 0 \ \sin \left[ \phi 
ight] & \cos \left[ \phi 
ight] & 0 \ 0 & 0 & 1 \end{array} egin{array}{c} 0 & 0 & 1 \end{array}$$

In[2]:= RotationMatrix[0, {0, 1, 0}] // MatrixForm

Out[2]//MatrixForm=

 $\begin{pmatrix} \cos\left[\theta\right] & 0 & \sin\left[\theta\right] \\ 0 & 1 & 0 \\ -\sin\left[\theta\right] & 0 & \cos\left[\theta\right] \end{pmatrix}$ 

In[3]:= RotationMatrix[a, {1, 0, 0}] // MatrixForm

Out[3]//MatrixForm=

by about 2 Rotat rotate by a about y (e; ten potate by & about 2" (?)  $\frac{4\pi i^2}{8\pi i^2} = \frac{8\pi i^2}{6\pi i^2} + \frac{6\pi i^2}{6\pi i^2} = \frac{6\pi i^2}{6\pi i^2} + \frac{6\pi$ 0 VIII COCOCY-SOSY - COSOCY-COSO 02 - 20 CØ SQ + SQ CQ - CO SQ SQ + CQCQ SQ SQ CO yin

Euler angles: can move (x, y, z) axes to arbitrary orientation.

#### In[2]:= RotationMatrix[\$\phi] // MatrixForm

#### Out[2]//MatrixForm=

 $\begin{pmatrix} \cos[\phi] & -\sin[\phi] \\ \sin[\phi] & \cos[\phi] \end{pmatrix}$ 

#### In[4]:= RotationMatrix[\$\phi\$, {0, 0, 1}] // MatrixForm

#### Out[4]//MatrixForm=

 $\begin{pmatrix} \cos\left[\phi\right] & -\sin\left[\phi\right] & 0\\ \sin\left[\phi\right] & \cos\left[\phi\right] & 0\\ 0 & 0 & 1 \end{pmatrix}$ 

#### In[5]:= RotationMatrix[0, {0, 1, 0}] // MatrixForm

#### Out[5]//MatrixForm=

```
\begin{pmatrix} \cos\left[\Theta\right] & 0 & \sin\left[\Theta\right] \\ 0 & 1 & 0 \\ -\sin\left[\Theta\right] & 0 & \cos\left[\Theta\right] \end{pmatrix}
\ln[10] = r1 = RotationMatrix[\phi, \{0, 0, 1\}];
r2 = RotationMatrix[\Theta, \{0, 1, 0\}];
r3 = RotationMatrix[\psi, \{0, 0, 1\}];
r3 , r2 , r1 // MatrixForm
```

Out[13]//MatrixForm=

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 \begin{pmatrix} \cos\left[\theta\right] \cos\left[\psi\right] - \sin\left[\phi\right] \sin\left[\psi\right] - \cos\left[\theta\right] \cos\left[\psi\right] \sin\left[\phi\right] - \cos\left[\theta\right] \sin\left[\psi\right] \cos\left[\psi\right] \sin\left[\psi\right] \cos\left[\psi\right] \sin\left[\theta\right] \\ \cos\left[\psi\right] \sin\left[\phi\right] + \cos\left[\theta\right] \cos\left[\phi\right] \sin\left[\psi\right] & \cos\left[\theta\right] \cos\left[\phi\right] - \cos\left[\theta\right] \sin\left[\phi\right] \sin\left[\psi\right] & \sin\left[\theta\right] \sin\left[\psi\right] \\ -\cos\left[\phi\right] \sin\left[\theta\right] & \sin\left[\theta\right] & \sin\left[\theta\right] & \cos\left[\theta\right] \sin\left[\phi\right] \\ \end{pmatrix} \end{cases}
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Suppose the Euler angles  $\phi$ ,  $\theta$ ,  $\psi$  vary with time, as the body orientation changes w.r.t. the space frame's fixed x, y, z axes. I'll write out more steps than Taylor does, and I may confuse you by saying  $(\hat{x}, \hat{y}, \hat{z}) \rightarrow (\hat{e}''_1, \hat{e}''_2, \hat{e}''_3) \rightarrow (\hat{e}'_1, \hat{e}'_2, \hat{e}'_3) \rightarrow (\hat{e}_1, \hat{e}_2, \hat{e}_3)$ . I do this so that my  $(\hat{e}'_1, \hat{e}'_2, \hat{e}'_3)$  are the same as Taylor's.

- 1. Rotate by  $\phi$  about  $\hat{m{z}} 
  ightarrow \hat{m{e}}_1''$ ,  $\hat{m{e}}_2''$ .  $(\hat{m{e}}_3''=\hat{m{z}}.)$
- 2. Rotate by heta about  $\hat{e}_2'' 
  ightarrow \hat{e}_1'$ ,  $\hat{e}_3'$ .  $(\hat{e}_2' = \hat{e}_2'')$ .
- 3. Rotate by  $\psi$  about  $\hat{e}'_3 
  ightarrow \hat{e}_1$ ,  $\hat{e}_2$ .  $(\hat{e}_3 = \hat{e}'_3.)$



Next, project  $\omega$  onto more convenient sets of unit vectors.

(orthogonal matrix: inverse = transpose)  $\hat{e}_{1}^{"} = \hat{\chi} \cos \phi + \hat{y} \sin \phi$  $\hat{e}_{2}^{"} = -\hat{\chi} \sin \phi + \hat{y} \cos \phi$  $\hat{e}_{3}^{"} = \hat{z}$  $\hat{x} = \hat{e}_{1}^{\prime\prime} (\cos \phi - \hat{e}_{2}^{\prime\prime} \sin \phi)$  $\hat{y} = \hat{e}_{1}^{\prime\prime} \sin \phi + \hat{e}_{2}^{\prime\prime} \cos \phi$  $\hat{e}_{1}^{\prime\prime} = \hat{e}_{1}^{\prime} (\cos + \hat{e}_{3}^{\prime} \sin \theta)$  $\hat{e}_{3}^{\prime\prime} = -\hat{e}_{1}^{\prime} \sin \theta + \hat{e}_{3}^{\prime} \cos \theta$  $\hat{e}'_{1} = \hat{e}''_{1} \cos \theta - \hat{e}''_{3} \sin \theta$   $\hat{e}'_{3} = \hat{e}''_{1} \sin \theta + \hat{e}''_{1} \cos \theta$   $\hat{e}'_{3} = \hat{e}''_{1} \sin \theta + \hat{e}''_{1} \cos \theta$  $\hat{\varrho}_1 = \hat{\varrho}_1' \cos \psi + \hat{\varrho}_2' \sin \psi$  $\hat{\varrho}_2 = -\hat{\varrho}_1' \sin \psi + \hat{\varrho}_1' \cos \psi$  $\hat{\varrho}_3 = \hat{\varrho}_3'$  $\hat{e}_1' = \hat{e}_1 \cos \psi - \hat{e}_2 \sin \psi$  $\hat{e}_2' = \hat{e}_1 \sin \psi + \hat{e}_2 \cos \psi$  $\omega = \dot{\phi} + \dot{\phi} + \dot{\phi} + \dot{\psi} + \dot{\phi} = \dot{\phi} + \dot{\phi$  $W = (-85in\# + 45in\# \cos \theta, 9\cos \# + 45in\# 5in\#, $$$ + 4000)$ (WSPACE AKES)

In the "space" basis [proof on previous page]:  $\boldsymbol{\omega} = (-\dot{\theta}\sin\phi + \dot{\psi}\sin\theta\cos\phi)\hat{\boldsymbol{x}} + (\dot{\theta}\cos\phi + \dot{\psi}\sin\theta\sin\phi)\hat{\boldsymbol{y}} + (\dot{\phi} + \dot{\psi}\cos\theta)\hat{\boldsymbol{z}}$ In the "body" basis [proof on next page]:

 $\boldsymbol{\omega} = (-\dot{\phi}\sin\theta\cos\psi + \dot{\theta}\sin\psi)\hat{\boldsymbol{e}}_1 + (\dot{\phi}\sin\theta\sin\psi + \dot{\theta}\cos\psi)\hat{\boldsymbol{e}}_2 + (\dot{\phi}\cos\theta + \dot{\psi})\hat{\boldsymbol{e}}_3$ 

Most convenient for symmetric top  $(\lambda_1 = \lambda_2)$ : in the "primed" basis (i.e. before the final rotation by  $\psi$  about  $\hat{e}_3$ ). Note that  $\hat{e}'_3 = \hat{e}_3$ .

$$\boldsymbol{\omega} = (-\dot{\phi}\sin\theta)\hat{\boldsymbol{e}}_1' + (\dot{\theta})\hat{\boldsymbol{e}}_2' + (\dot{\phi}\cos\theta + \dot{\psi})\hat{\boldsymbol{e}}_3'$$

This last one is easiest to see if you consider the instant at which  $\psi = 0$ .

 $\hat{e}_{,''=} + \hat{y} \cos \phi + \hat{y} \sin \phi$  $\hat{e}_{,''=} - \hat{x} \sin \phi + \hat{y} \cos \phi$  $\hat{e}_{,''=} = \hat{z}$  $\hat{x} = \hat{e}_{1}^{"}(\cos\phi - \hat{e}_{2}^{"}\sin\phi)$  $\hat{y} = \hat{e}_{1}^{\prime\prime} \sin x + \hat{e}_{2}^{\prime\prime} \cos x$  $\hat{e}'_{1} = \hat{e}''_{1} \cos \theta - \hat{e}''_{3} \sin \theta \\
\hat{e}'_{3} = \hat{e}''_{1} \sin \theta + \hat{e}''_{3} \cos \theta \\
\hat{e}'_{2} = \hat{e}''_{3} \sin \theta + \hat{e}''_{3} \cos \theta \\
\hat{e}'_{2} = \hat{e}''_{3} \sin \theta \\
\hat{e}'_{2} = \hat{e}''_{3} \sin \theta \\
\hat{e}'_{3} = \hat{e}''_{3} \sin^{2} \theta \\
\hat{e}''_{3} = \hat{e}''_{3} \sin^{2} \theta \\
\hat{e}''_{3$  $\hat{e}_{3}'' = \hat{e}_{3}' (so + \hat{e}_{3}'s) \hat{n} o$  $\hat{e}_{3}'' = -\hat{e}_{3}'s \hat{n} o + \hat{e}_{3}' coro$  $\hat{e}_1 = \hat{e}_1 \cos \psi + \hat{e}_2 \sin \psi$  $\hat{e}'_{1} = \hat{e}_{1}\cos\psi - \hat{e}_{2}\sin\psi$  $\hat{e}_{2}' = \hat{e}_{1}\sin\psi + \hat{e}_{2}\cos\psi$  $= -\hat{e}_{1}^{\prime} \sin \psi + \hat{e}_{2}^{\prime} \cos \psi$  $\hat{e}_{2} = \hat{e}_{2}^{\prime}$  $\omega = \phi \hat{z} + \theta \hat{e}_2'' + \psi \hat{e}_3''$  $= \phi e_3' + \partial e_2'' + \gamma e_3'$  $= \phi(-\hat{e}_1' \sin \theta + \hat{e}_3' \cos \theta) + \hat{\theta}(\hat{e}_2') + \hat{\psi}(\hat{e}_3')$  $-\dot{\phi}sin\phi(\hat{e}_{1}')+\dot{\phi}cos\phi(\hat{e}_{3}')+\dot{\phi}(\hat{e}_{2}')+\dot{\psi}(\hat{e}_{3}')$ =  $[-\phi\sin\theta(\hat{e}_1sy - \hat{e}_2sin\psi) + \phi\cos\theta\hat{e}_3] + \dot{\theta}(\hat{e}_1sin\psi + \hat{e}_2\cos\psi) + \dot{\psi}\hat{e}_3'$  $= \hat{e}_1 \left( -\dot{\phi} \sin \theta \cos \psi + \dot{\theta} \sin \psi \right) + \hat{e}_2 \left( \dot{\phi} \sin \theta \sin \psi + \dot{\theta} \cos \psi \right) + \hat{e}_3 \left( \dot{\phi} \cos \theta + \dot{\psi} \right)$ 

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Most convenient for symmetric top  $(\lambda_1 = \lambda_2)$ :

$$\boldsymbol{\omega} = (-\dot{\phi}\sin\theta)\hat{\boldsymbol{e}}_1' + (\dot{\theta})\hat{\boldsymbol{e}}_2' + (\dot{\phi}\cos\theta + \dot{\psi})\hat{\boldsymbol{e}}_3'$$

Now, how do we write the top's angular momentum L and kinetic energy T? How about the Lagrangian?

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Most convenient for symmetric top  $(\lambda_1 = \lambda_2)$ :

$$\boldsymbol{\omega} = (-\dot{\phi}\sin\theta)\hat{\boldsymbol{e}}_1' + (\dot{\theta})\hat{\boldsymbol{e}}_2' + (\dot{\phi}\cos\theta + \dot{\psi})\hat{\boldsymbol{e}}_3'$$

Now, how do we write the top's angular momentum L and kinetic energy T? How about the Lagrangian?

 $\boldsymbol{L} = (-\lambda_1 \dot{\phi} \sin \theta) \boldsymbol{\hat{e}}_1' + (\lambda_1 \dot{\theta}) \boldsymbol{\hat{e}}_2' + \lambda_3 (\dot{\phi} \cos \theta + \dot{\psi}) \boldsymbol{\hat{e}}_3'$ 

$$T = \frac{1}{2}\lambda_1(\dot{\phi}^2\sin^2\theta + \dot{\theta}^2) + \frac{1}{2}\lambda_3(\dot{\phi}\cos\theta + \dot{\psi})^2$$

$$\mathcal{L} = \frac{1}{2}\lambda_1(\dot{\phi}^2\sin^2\theta + \dot{\theta}^2) + \frac{1}{2}\lambda_3(\dot{\phi}\cos\theta + \dot{\psi})^2 - MgR\cos\theta$$

Which coordinates (Euler angles) are ignorable? What are the corresponding conserved momenta?

$$\boldsymbol{\omega} = (-\dot{\phi}\sin\theta)\hat{\boldsymbol{e}}_1' + (\dot{\theta})\hat{\boldsymbol{e}}_2' + (\dot{\phi}\cos\theta + \dot{\psi})\hat{\boldsymbol{e}}_3'$$
$$\boldsymbol{L} = (-\lambda_1\dot{\phi}\sin\theta)\hat{\boldsymbol{e}}_1' + (\lambda_1\dot{\theta})\hat{\boldsymbol{e}}_2' + \lambda_3(\dot{\phi}\cos\theta + \dot{\psi})\hat{\boldsymbol{e}}_3'$$
$$\boldsymbol{\mathcal{L}} = \frac{1}{2}\lambda_1(\dot{\phi}^2\sin^2\theta + \dot{\theta}^2) + \frac{1}{2}\lambda_3(\dot{\phi}\cos\theta + \dot{\psi})^2 - MgR\cos\theta$$

$$\frac{\partial \lambda}{\partial \dot{\psi}} = \lambda_3 (\dot{\psi} + \dot{\phi} \cos \theta) = P_{\psi} = \cos \theta + 2_3 \omega_3 = 2_3$$

$$\frac{\partial \chi}{\partial \dot{\psi}} = \lambda_3 (\dot{\psi} + \dot{\phi} \cos \theta) = P_{\psi} = \cos \theta$$

$$\frac{\partial \chi}{\partial \dot{\psi}} = \lambda_3 (\dot{\psi} + \dot{\phi} \cos \theta) \cos \theta = P_{\psi} = \cos \theta$$

$$= (L_2 - L_3 \cos \theta) + L_3 \cos \theta = 2_2$$

$$Digrassion: \hat{z} = -\hat{e}_1' \sin \theta + \hat{e}_3' \cos \theta$$

$$L_2 = 2 \cdot \hat{z} = \lambda_3 \dot{\phi} \sin^2 \theta + \lambda_3 (\dot{\psi} + \dot{\phi} \cos \theta) \cos \theta$$

$$L_2 = 2 \cdot \hat{z} = \lambda_3 \dot{\phi} \sin^2 \theta + 2_3 \cos \theta$$

$$\Rightarrow \lambda_3 \dot{\phi} \sin^2 \theta = L_2 - L_3 \cos \theta$$

$$\mathcal{L} = \frac{1}{2}\lambda_{1}(\dot{\phi}^{2}\sin^{2}\theta + \dot{\theta}^{2}) + \frac{1}{2}\lambda_{3}(\dot{\phi}\cos\theta + \dot{\psi})^{2} - MgR\cos\theta$$

$$O = \frac{1}{2}\lambda_{1}(\dot{\phi}^{2}\sin^{2}\theta + \dot{\theta}^{2}) + \frac{1}{2}\lambda_{3}(\dot{\phi}\cos\theta + \dot{\psi})^{2} - MgR\cos\theta$$

$$L_{1}\dot{\phi} = \lambda_{1}(\dot{\phi}^{2}\sin^{2}\theta + M_{2}R_{1})\phi$$

$$L_{1}\dot{\phi} = \lambda_{1}(\dot{\phi}^{2}\sin^{2}\theta + M_{2}R_{2})\phi$$
First consider case where  $\vartheta = constant$ . Since  $\lambda_{1}\dot{\phi}\sin^{2}\theta = L_{2} - L_{3}\cos\theta$ ,  $\vartheta = constant$ . Since  $\lambda_{1}\dot{\phi}\sin^{2}\theta = L_{2} - L_{3}\cos\theta$ ,  $\vartheta = const \Rightarrow \dot{\phi} = const = 1$ 

$$O = \lambda_{1}(D_{2}^{2}\sin^{2}(\cos\theta - \lambda_{2}\omega_{2})R_{2})\phi + M_{2}R_{2}i\theta$$

$$(\lambda_{1}\cos^{2}\theta)R_{2}^{2} - (\lambda_{3}\omega_{2})R_{2} + M_{2}R_{2} = 0$$

$$R = \frac{\lambda_{2}\omega_{2} \pm (\lambda_{2}\omega_{2})R_{2} + M_{2}R_{2}}{(\lambda_{3}\omega_{3})^{2}}$$

$$R = \frac{\lambda_{3}\omega_{3}}{(\lambda_{3}\cos^{2}\theta)}(1 \pm (1 - \frac{4\lambda_{1}\cos^{2}\theta}{(\lambda_{3}\omega_{3})^{2}})$$

 $\mathcal{T} = \frac{\lambda_3 \omega_3}{2\lambda_1 \cos \theta} \left( \left| \frac{1}{2} \right| \right) - \frac{4\lambda_1 \cos \theta M_2 R}{(\lambda_3 \omega_3)^2}$ has 2 real solutions if (13w3) > 41, MgRcord ("if wy is large enough") Math is simplest if  $(\lambda_z \omega_z)^2 \gg 4\lambda_1 M_9 Reprod$ ("wz is vory large") $=\frac{2\lambda_1\cos^2}{\lambda_3\omega_3}$ (Jzwz)2  $\int \sum_{n=1}^{\infty} \frac{\lambda_3 \omega_3}{2\lambda_1 (\omega_3)^2} \frac{2\lambda_1 M_3 R(\omega_3)}{(\lambda_3 \omega_3)^2} =$  $\simeq \frac{\lambda_3 \omega_3}{\lambda_3 \omega_3}$ A, coro This can be seen from with no torque see HP GNIZS \_WER = GNIS

(Skip: Just in case you wanted to see the  $\theta$  EOM derived.)

$$\mathcal{L} = \frac{1}{2}\lambda_1(\dot{\phi}^2\sin^2\theta + \dot{\theta}^2) + \frac{1}{2}\lambda_3(\dot{\phi}\cos\theta + \dot{\psi})^2 - MgR\cos\theta$$

$$\frac{\partial \mathcal{G}}{\partial \theta} = \lambda_{i}\dot{\phi}^{2}\sin\theta\cos\theta + \lambda_{3}(\dot{\phi}\cos\theta+\dot{\phi})(-\dot{\phi}\sin\theta) + MgRsin\theta}$$

$$\frac{d}{\partial \theta}(\frac{\partial \mathcal{G}}{\partial \theta}) = \frac{d}{\partial \theta}(\lambda_{i}\dot{\theta}) = \lambda_{i}\dot{\theta}^{2} = \lambda_{i}\dot{\theta}^{$$

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 $\lambda_1 \phi \sin^2 \theta = L_2 - L_3 \cos \theta \Rightarrow \phi = \frac{L_2 - L_3 \cos \theta}{\lambda_1 \sin^2 \theta}$  $E = T + U = \frac{1}{2}\lambda_1 \left(\phi^2 \sin^2 \theta + \phi^2\right) + \frac{1}{2}\lambda_2 \left(\psi + \phi \cos \theta^2 + M_g R_{COS} \theta + M_g$  $E = \frac{1}{2}\lambda_1 \sin^2 \theta \dot{\theta}^2 + \frac{1}{2}\lambda_1 \dot{\theta}^2 + \frac{1}{2}\lambda_3 w_3^2 + M_9 R_{COTO}$  $E = \frac{\lambda_{sl_{n}^{29}} (L_{2} - L_{3} \cos \theta)^{2}}{2} + \frac{\lambda_{10}^{2}}{\lambda_{1}^{2} \sin^{9} \theta} + \frac{\lambda_{10}^{2}}{2} + \frac{L_{3}^{2}}{2\lambda_{3}} + M_{9}Rcos\theta$  $E = \frac{\lambda_1 \delta^2}{2} + \frac{(L_2 - L_3 \cos)^2}{2\lambda_1 \sin^2 \Theta} + \frac{L_3^2}{2\lambda_3} + M_3 R \cos \Theta$ Ueff (0)  $E = \pm \lambda_1 \theta + N_{eff}(\theta) \quad (one timensional problem)$   $\theta = \frac{1}{2}\lambda_1 \theta + N_{eff}(\theta) \quad (one timensional problem)$ 

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From the final exam for the course I took, fall 1990. (This turns out to be the same problem as appears in Feynman's story of the cafeteria plate that wobbles as it flies through the air.)

An infinitely thin, uniform, square plate of mass m and side d is allowed to undergo rotation. At time t = 0, the normal to the plate,  $\hat{e}_3$ , is aligned with  $\hat{z}$ , but the angular velocity vector  $\underline{\omega}$  deviates from  $\hat{z}$  by a small angle  $\alpha$ . Work the entire problem to first order in  $\alpha$ , i.e. drop terms of  $0(\alpha^2)$  or higher.



(a) Show I =  $I_0 \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 2 \end{pmatrix}$  and find  $I_0$ .

- (b) Find the maximum angle between  $\widehat{z}$  and  $\widehat{e_3}$  during subsequent motion of the plate.
- (c) When is this maximum deviation first reached?

J= then t, w, = 1;  $I_{X,=Z_{r}}$  $I \in \lambda$ If  $\Gamma_3 = 0$  then  $\omega_3 = const.$  Suppose  $\Gamma_3 = 0.$  $\omega_{1} = P_{1} + (\lambda - \lambda_{3}) \omega_{2} \omega_{3}$ let (23-2) WZW  $\dot{w}_2 = P_2 +$  $W_1 = \mathcal{N}_W_2$ ,  $W_2 = -\mathcal{N}_W_2$ -> W = W cosset No =- No shilt E W2

As seen from body frame, L and  $\omega$  precess about (fixed)  $\hat{e}_3$  with frequency  $\Omega_b \equiv \Omega = \omega_3(\lambda - \lambda_3)/\lambda$ , where  $\lambda = \lambda_1 = \lambda_2$ .

As seen from the space frame,  $\hat{e}_3$  and  $\omega$  precess about (fixed) L, at frequency  $\Omega_s = L/\lambda_1$ , which you'll prove in the HW.

8. [This problem is adapted from a problem on the final exam I took for the analogous course in fall 1990.] A uniform, infinitesimally thick, square plate of mass m and side length d is allowed to undergo torque-free rotation. At time t = 0, the normal to the plate,  $\hat{\boldsymbol{e}}_3$ , is aligned with  $\hat{\boldsymbol{z}}$ , but the angular velocity vector  $\boldsymbol{\omega}$  deviates from  $\hat{\boldsymbol{z}}$  by a small angle  $\alpha$ . The figure below depicts the situation at time t = 0, at which time  $\hat{\boldsymbol{e}}_1 = \hat{\boldsymbol{x}}$ ,  $\hat{\boldsymbol{e}}_2 = \hat{\boldsymbol{y}}$ ,  $\hat{\boldsymbol{e}}_3 = \hat{\boldsymbol{z}}$ , and  $\boldsymbol{\omega} = \boldsymbol{\omega}(\cos \alpha \hat{\boldsymbol{z}} + \sin \alpha \hat{\boldsymbol{x}})$ .



time t=0 (a) Show that  $\underline{I} = I_0 \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 2 \end{pmatrix}$ and find the constant  $I_0$ .

(b) Calculate  $\boldsymbol{L}$  at t = 0.

 $\hat{x}$  (c) Sketch  $\hat{e}_3$ ,  $\omega$ , and L at t=0.

(d) Draw/label "body cone" and "space cone" on your sketch.

8. [This problem is adapted from a problem on the final exam I took for the analogous course in fall 1990.] A uniform, infinitesimally thick, square plate of mass m and side length d is allowed to undergo torque-free rotation. At time t = 0, the normal to the plate,  $\hat{\boldsymbol{e}}_3$ , is aligned with  $\hat{\boldsymbol{z}}$ , but the angular velocity vector  $\boldsymbol{\omega}$  deviates from  $\hat{\boldsymbol{z}}$  by a small angle  $\alpha$ . The figure below depicts the situation at time t = 0, at which time  $\hat{\boldsymbol{e}}_1 = \hat{\boldsymbol{x}}$ ,  $\hat{\boldsymbol{e}}_2 = \hat{\boldsymbol{y}}$ ,  $\hat{\boldsymbol{e}}_3 = \hat{\boldsymbol{z}}$ , and  $\boldsymbol{\omega} = \boldsymbol{\omega}(\cos \alpha \hat{\boldsymbol{z}} + \sin \alpha \hat{\boldsymbol{x}})$ .



(e) Calculate precession frequencies  $\Omega_{body}$  and  $\Omega_{space}$ . Indicate directions of precession vectors  $\Omega_{body}$  and  $\Omega_{space}$  on drawing.

(f) You argue in HW that  $\Omega_{space} = \Omega_{body} + \omega$ . Verify (by writing out components) that this relationship holds for the  $\Omega_{space}$  and  $\Omega_{body}$  that you calculate for t = 0. 8. [This problem is adapted from a problem on the final exam I took for the analogous course in fall 1990.] A uniform, infinitesimally thick, square plate of mass m and side length d is allowed to undergo torque-free rotation. At time t = 0, the normal to the plate,  $\hat{\boldsymbol{e}}_3$ , is aligned with  $\hat{\boldsymbol{z}}$ , but the angular velocity vector  $\boldsymbol{\omega}$  deviates from  $\hat{\boldsymbol{z}}$  by a small angle  $\alpha$ . The figure below depicts the situation at time t = 0, at which time  $\hat{\boldsymbol{e}}_1 = \hat{\boldsymbol{x}}$ ,  $\hat{\boldsymbol{e}}_2 = \hat{\boldsymbol{y}}$ ,  $\hat{\boldsymbol{e}}_3 = \hat{\boldsymbol{z}}$ , and  $\boldsymbol{\omega} = \boldsymbol{\omega}(\cos \alpha \hat{\boldsymbol{z}} + \sin \alpha \hat{\boldsymbol{x}})$ .



(g) Find the maximum angle between  $\hat{z}$  and  $\hat{e}_3$  during subsequent motion of the plate. Show that in the limit  $\alpha \ll 1$ , this maximum angle equals  $\alpha$ .

(h) When is this maximum deviation first reached?

video: https://www.youtube.com/
watch?v=oH-dlrIF010

# Physics 351 — Wednesday, March 28, 2018

- I have graded 40/60 of the midterm exams. I'll finish tomorrow and will hand them back Friday.
- HW9 due Friday. For HW help, Bill is in DRL 3N6 Wed 4–7pm. Grace is in DRL 2C2 Thu 5:30–8:30pm. To get the most benefit from the homework, first work through every problem on your own to the best of your ability. Then check in with me, Grace, or a friend to compare final results and to trade suggestions on problems that stumped you.
- For the next 3–4 days, we will finally make our way through the rest of Chapter 10, which is (I think) the most difficult chapter in the book. The two significant topics after that are coupled oscillators and the Hamiltonian formalism, plus a bit of "enrichment" material.
- ► FYI intuitive description of precession:

http://positron.hep.upenn.edu/p351/files/0331\_george\_abell\_precession.pdf