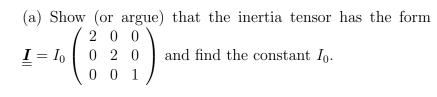
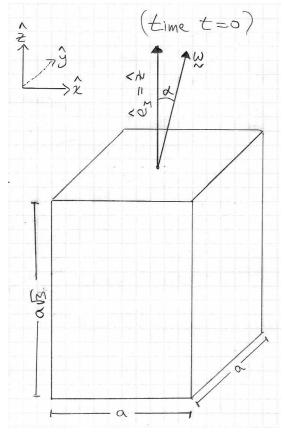
## Problem 1.

A uniform rectangular solid of mass m and dimensions  $a \times a \times a\sqrt{3}$  (volume  $\sqrt{3} a^3$ ) is allowed to undergo torque-free rotation. At time t = 0, the long axis (length  $a\sqrt{3}$ ) of the solid is aligned with  $\hat{z}$ , but the angular velocity vector  $\boldsymbol{\omega}$  deviates from  $\hat{z}$  by a small angle  $\alpha$ . The figure depicts the situation at time t = 0, at which time  $\hat{\boldsymbol{e}}_1 = \hat{\boldsymbol{x}}$ ,  $\hat{\boldsymbol{e}}_2 = \hat{\boldsymbol{y}}$ ,  $\hat{\boldsymbol{e}}_3 = \hat{\boldsymbol{z}}$ , and  $\boldsymbol{\omega} = \omega(\cos\alpha\hat{\boldsymbol{z}} + \sin\alpha\hat{\boldsymbol{x}})$ .





(b) Calculate the angular momentum vector  $\boldsymbol{L}$  at t = 0. Write  $\boldsymbol{L}(t = 0)$  both in terms of  $\hat{\boldsymbol{e}}_1, \hat{\boldsymbol{e}}_2, \hat{\boldsymbol{e}}_3$  and in terms of  $\hat{\boldsymbol{x}}, \hat{\boldsymbol{y}}, \hat{\boldsymbol{z}}$ . Which of these two expressions will continue to be valid into the future?

(c) Draw a sketch showing the vectors  $\hat{\boldsymbol{e}}_3$ ,  $\boldsymbol{\omega}$ , and  $\boldsymbol{L}$  at t = 0. Be sure that the relative orientation of  $\boldsymbol{L}$  and  $\boldsymbol{\omega}$  makes sense. This relative orientation is different for egg-shaped ("prolate") objects  $(\lambda_3 < \lambda_1)$  than it is for frisbee-like ("oblate") objects  $(\lambda_3 > \lambda_1)$ .

(d) Draw and label the "body cone" and the "space cone" on your sketch.

(e) Calculate the precession frequencies  $\Omega_{\text{body}}$  and  $\Omega_{\text{space}}$ . Indicate the directions of the precession vectors  $\Omega_{\text{body}}$  and  $\Omega_{\text{space}}$  on your drawing. Be careful with the "sign" of the  $\Omega_{\text{body}}$  vector, i.e. be careful not to draw  $-\Omega_{\text{body}}$  when you mean to draw  $\Omega_{\text{body}}$ .

(f) You argued in HW11 that  $\Omega_{\text{space}} = \Omega_{\text{body}} + \omega$ . Verify (by writing out components) that this relationship holds for the  $\Omega_{\text{space}}$  and  $\Omega_{\text{body}}$  that you calculate for t = 0.

(g) In the  $\alpha \ll 1$  limit (so  $\tan \alpha \approx \alpha$ ,  $\tan(2\alpha) \approx 2\alpha$ , etc.), find the maximum angle between  $\hat{z}$  and  $\hat{e}_3$  during subsequent motion of the solid. (This should be some constant factor times  $\alpha$ .) A simple argument is sufficient here, no calculation.

(h) At what time t is this maximum deviation first reached?

(This problem shows that for an American-football-like object, the frequency of the wobbling motion is smaller than the frequency of the spinning motion — which is opposite the conclusion that you reached for the flying dinner plate, whose wobbling was twice as fast as its spinning.) Possibly useful equations.

$$\left(\frac{\mathrm{d}\boldsymbol{Q}}{\mathrm{d}t}\right)_{\mathrm{space}} \;\; = \;\; \left(\frac{\mathrm{d}\boldsymbol{Q}}{\mathrm{d}t}\right)_{\mathrm{body}} \;\; + \;\; \boldsymbol{\Omega} imes \boldsymbol{Q}$$

$$m\ddot{\boldsymbol{r}} = \boldsymbol{F} + 2m\,\dot{\boldsymbol{r}} imes \boldsymbol{\Omega} + m\,(\boldsymbol{\Omega} imes \boldsymbol{r}) imes \boldsymbol{\Omega} = \boldsymbol{F} + 2m\,\boldsymbol{v} imes \boldsymbol{\Omega} + m\Omega^2 
ho\,\hat{\boldsymbol{
ho}}$$

For a uniform solid cylinder of radius R about its symmetry axis,  $I = mR^2/2$ . For a uniform thin rod of length L about its center (perpendicular to the rod axis),  $I = mL^2/12$ . For a rectangular plate about its center (rotation axis normal to plate),  $I = m(a^2 + b^2)/12$ , where a and b are the short and long side lengths.

For a free symmetric top,  $\Omega_s = L/\lambda_1$ 

Euler equations:

$$\tau_1 = \lambda_1 \dot{\omega_1} - (\lambda_2 - \lambda_3) \omega_2 \omega_3$$
  
$$\tau_2 = \lambda_2 \dot{\omega_2} - (\lambda_3 - \lambda_1) \omega_3 \omega_1$$
  
$$\tau_3 = \lambda_3 \dot{\omega_3} - (\lambda_1 - \lambda_2) \omega_1 \omega_2$$

If  $\boldsymbol{\tau} = 0$  and  $\lambda_1 = \lambda_2$  then the Euler equations reduce to the simpler form

$$\dot{\omega_3} = \frac{\lambda_1 - \lambda_1}{\lambda_3} \,\omega_1 \omega_2 = 0$$

$$\dot{\omega_1} = \frac{\lambda_1 - \lambda_3}{\lambda_1} \omega_2 \omega_3 = -\left(\frac{\lambda_3 - \lambda_1}{\lambda_1} \omega_3\right) \omega_2 = -\Omega_b \omega_2$$
$$\dot{\omega_2} = \frac{\lambda_3 - \lambda_1}{\lambda_1} \omega_3 \omega_1 = \left(\frac{\lambda_3 - \lambda_1}{\lambda_1} \omega_3\right) \omega_1 = \Omega_b \omega_1$$

so we can represent the precession of the  $\omega$  vector as an angular velocity vector  $\Omega_b$  with

$$\mathbf{\Omega}_b = \left(rac{\lambda_3 - \lambda_1}{\lambda_1}\,\omega_3
ight)\, \hat{oldsymbol{e}}_3.$$

Euler-angle convention: Start with body axes aligned with space axes. (i) Rotate body through angle  $\phi$  about  $\hat{z}$ . This leaves  $\hat{e}_3$  alone but rotates the first and second body axes in the xy plane. In particular, the second body axis now points in a direction called  $\hat{e}'_2$ . (ii) Rotate body through angle  $\theta$  about the new axis  $\hat{e}'_2$ . This moves the body axis  $\hat{e}_3$  to the direction whose polar angles are  $\theta$  and  $\phi$ . (iii) Rotate the body about  $\hat{e}_3$  through whatever angle  $\psi$  is needed to bring the body axes  $\hat{e}_2$  and  $\hat{e}_1$  into their assigned directions.

At any instant, you can use the values of  $\phi, \theta, \psi$  at that instant to write each body unit vector  $\hat{\boldsymbol{e}}_i$  as a linear combination of  $\hat{\boldsymbol{x}}, \hat{\boldsymbol{y}}, \hat{\boldsymbol{z}}$ . The coefficients involve sines and cosines of  $\phi, \theta, \psi$  but have no explicit time dependence.